

Development of fiber-based laser systems for LISA

Kenji Numata, Jordan Camp

Abstract

We present efforts on fiber-based laser systems for the LISA mission at the NASA Goddard Space Flight Center. A fiber-based system has the advantage of higher robustness against external disturbances and easier implementation of redundancies. For a master oscillator, we are developing a ring fiber laser and evaluating two commercial products, a DBR linear fiber laser and a planar-waveguide external cavity diode laser. They all have comparable performance to a traditional NPRO at LISA band. We are also performing reliability tests of a 2-W Yb fiber amplifier and radiation tests of fiber laser/amplifier components. We describe our progress to date and discuss the path to a working LISA laser system design.



Beyond Einstein: From the Big Bang to Black Holes

Development of fiber-based laser systems for LISA

Kenji Numata¹⁾²⁾, Jordan Camp²⁾

¹⁾Department of Astronomy, University of Maryland, CRESST

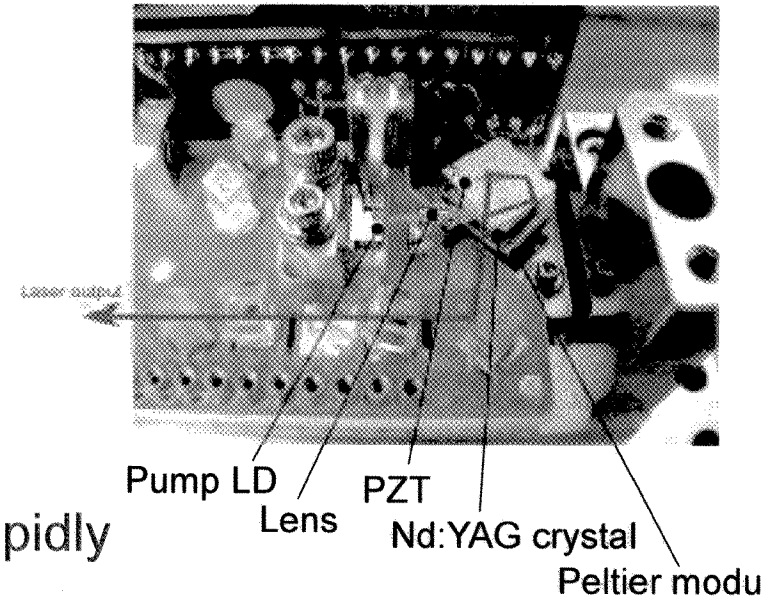
²⁾Gravitational Astrophysics Laboratory (Code 663), NASA/GSFC

- 1. Introduction
 - Motivation of this activity
- 2. Fiber-based lasers
 - GSFC ring fiber laser
 - NP photonics DBR fiber laser
 - RIO planar-waveguide external cavity diode laser
- 3. Fiber amplifier
- 4. Other activities
 - Space qualification tests
 - Fiber-based frequency stabilization
- 5. Summary

1. Introduction

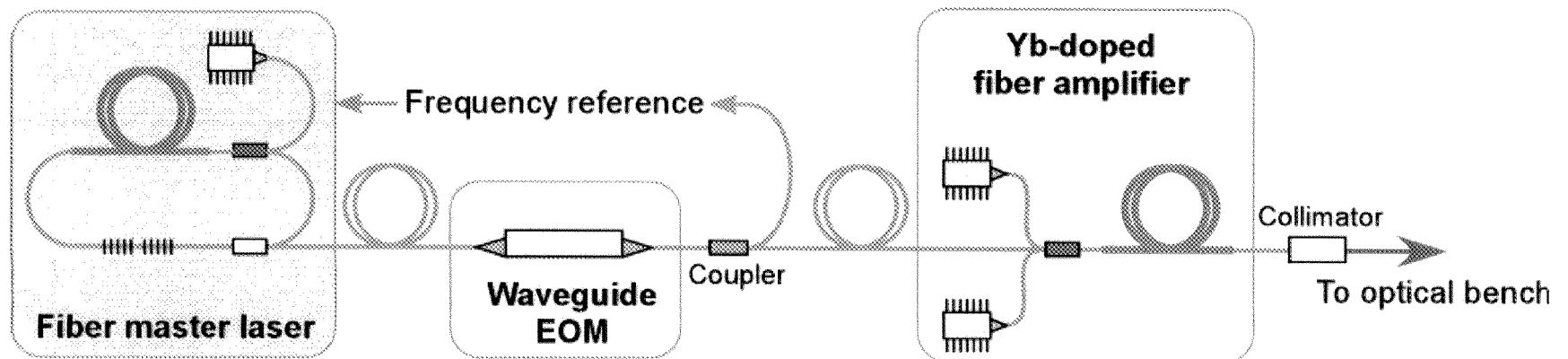
● NPRO (Non-planar ring oscillator) has been used traditionally.

- Compact crystal cavity gives high stability.
- “Black box” in many cases
 - E.g.) TESAT NPRO for LPF



● All fiber/waveguide solution

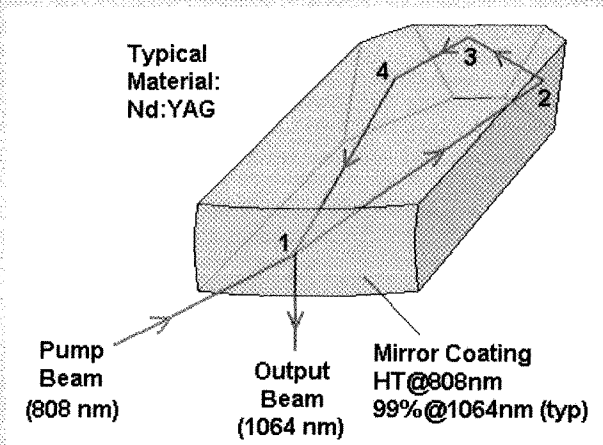
- Fiber laser/amplifier technologies matured rapidly
- Higher robustness



NPRO vs fiber laser

- Fiber laser offers significant advantages over NPRO laser

Traditional: NPRO laser



Difficult alignment

Glue/solder needed

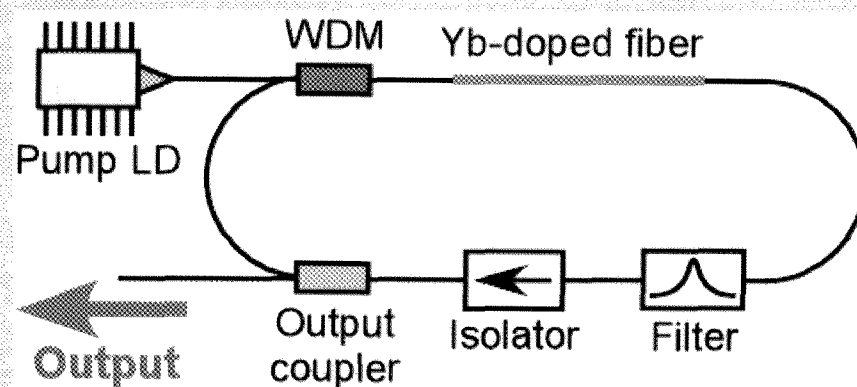
Need to couple back into fiber

Strong magnet needed

Contamination sensitive (sealed package)

Distorted Gaussian beam

New: Fiber laser



No alignment needed

No glue needed

Laser light within fiber

No strong magnet

No contamination

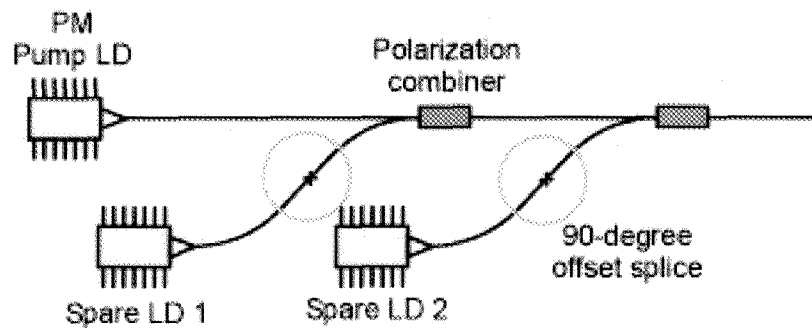
Mode & polarization cleaned by fiber

Fiber amp has..

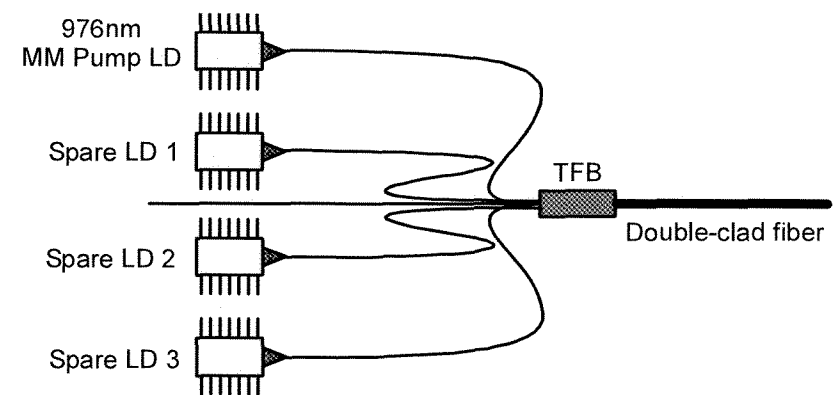
- Higher beam quality, lower sensitivity to alignment etc., easier cooling
- Higher reliability, **higher optical/wallplug efficiency**
 - E.g. Commercial fiber amp: >10% wallplug efficiency
 - ~2% efficiency in solid state amps in flight missions

Easier addition of redundancy

- Many (~90%?) laser failures come from pump LD
- No geometrical constraints



■ Core pumping by SM LDs



■ Clad pumping by MM LDs

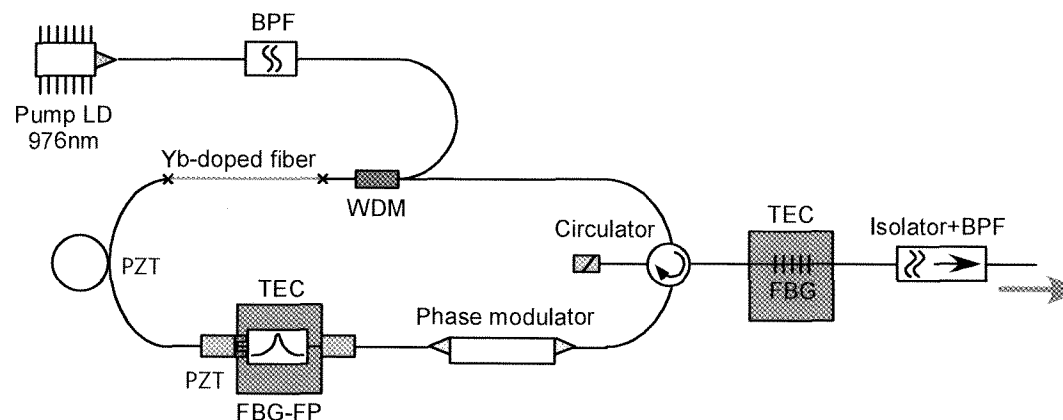
- GSFC fiber ring laser
 - Commercial highly-doped gain fiber + fiber Bragg gratings

- NP photonics/Fibertek DBR fiber laser
 - Special phosphosilicate glass fiber + fiber Bragg gratings

- RIO external cavity diode laser (ECL)
 - InP semiconductor gain chip + planar-waveguide Bragg reflector

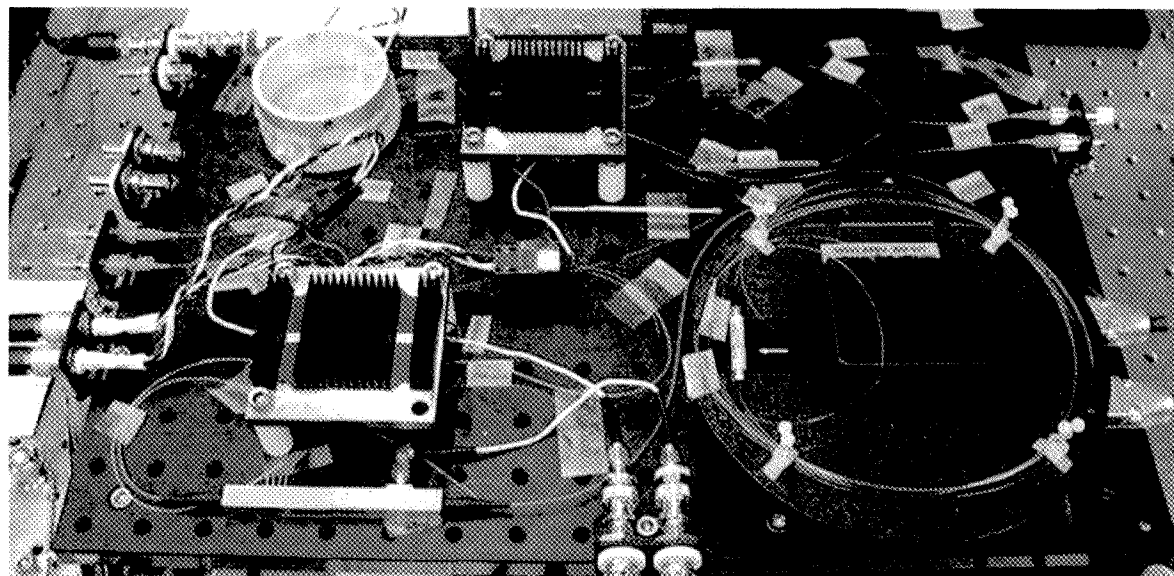
Features

- Design & built in house
- Commercial components only
 - No special gain fiber
 - No patent issues
- Two FBGs for single-mode selection
- Fast frequency tuning by waveguide EOM
- Low power ($\sim 2\text{mW}$)

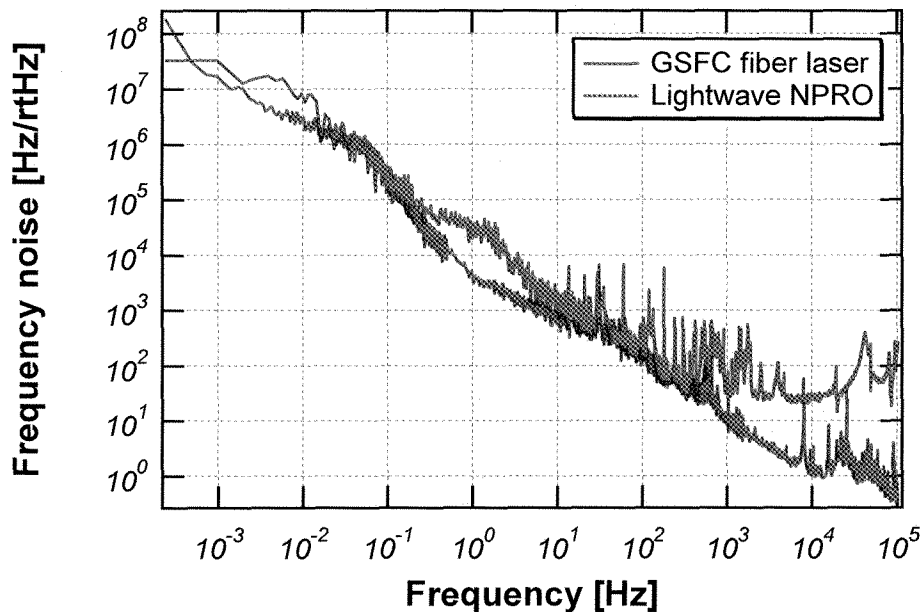


Status

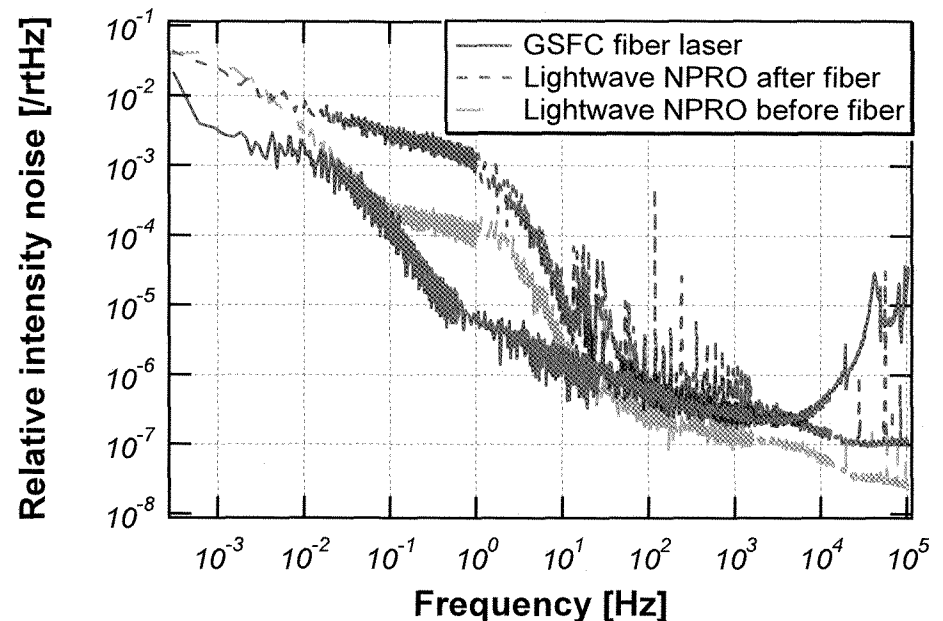
- Design fixed
- Iodine stabilization
- Digital system design



Frequency noise



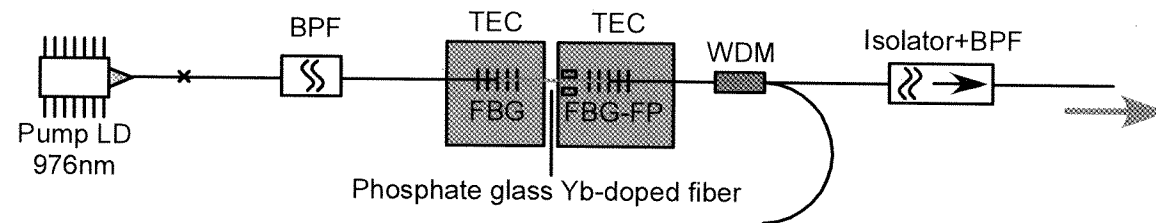
Intensity noise



- Low frequency: comparable to (better than) NPRO
- High frequency: increased noise due to relaxation oscillation
- Stabilization experiments
 - Frequency: Planned using iodine or cavity.
 - Intensity: Done after Yb amplifier and satisfied LISA requirement at low frequency.

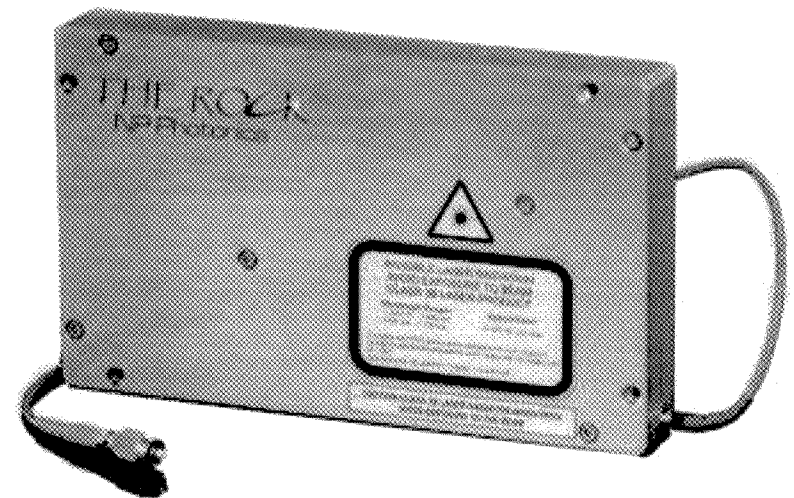
Features

- Built by NP Photonics
- Highly-doped phosphate glass fiber
 - Short cavity length
- Low reliability of splice
- Patented



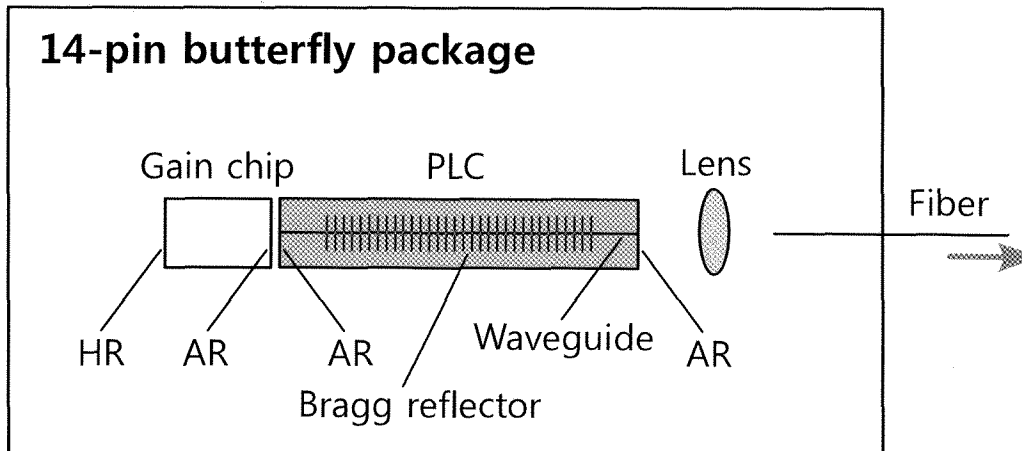
Status

- Qualification tests by Fibertek
 - “Space version” passed thermal cycling
- Noise evaluations



Features

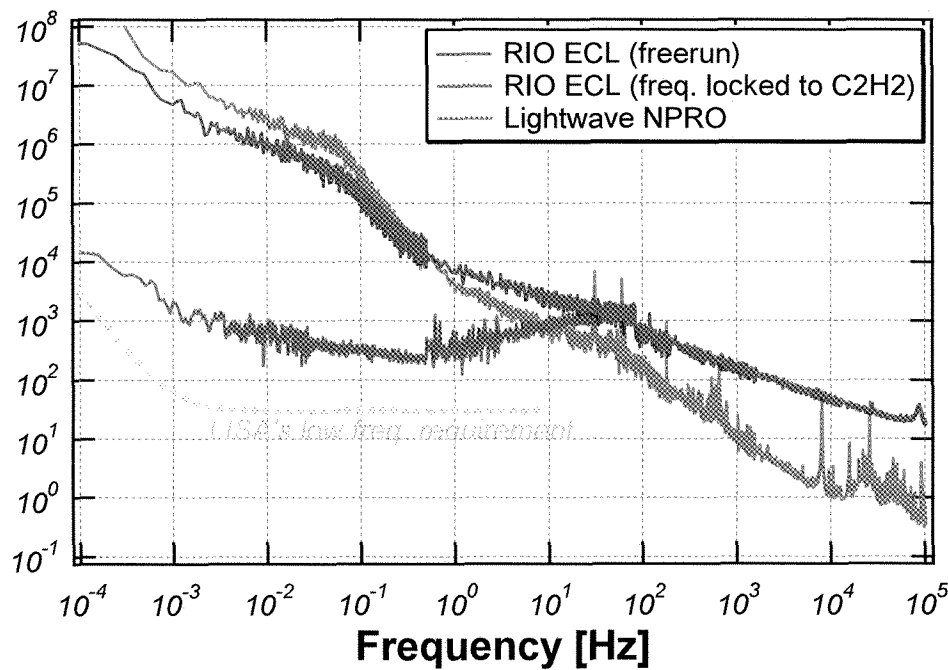
- Compact & simple
- **Low cost**
 - ~\$5k
- **Lowest noise** at LISA band
- Unconditionally single-mode
- Low power (~15mW)
- Telecom C-band only



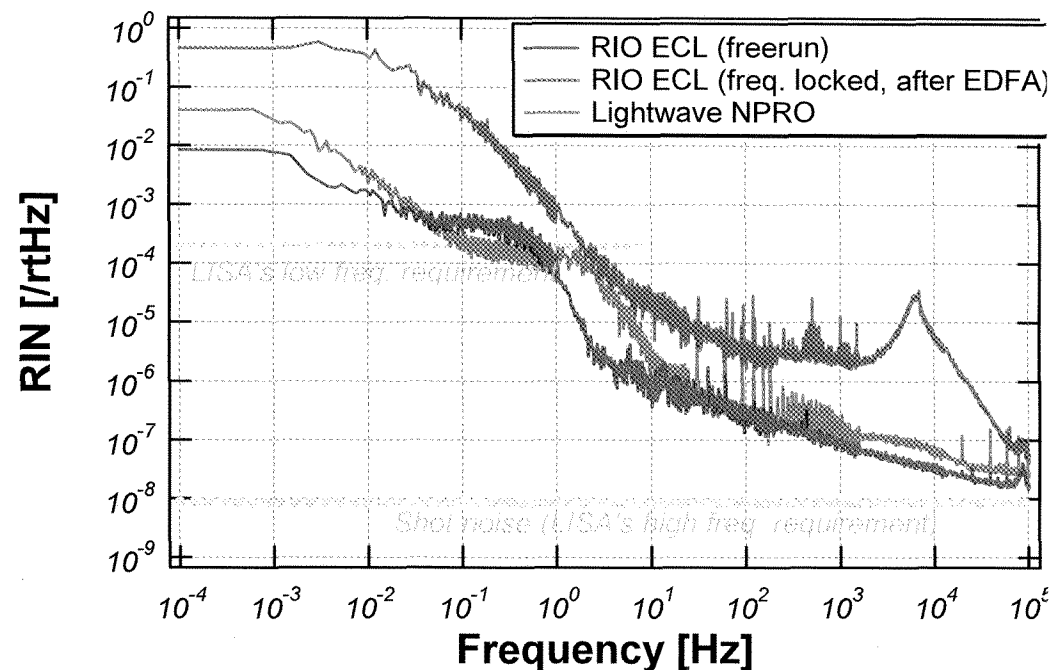
Status

- Frequency stabilization done
- Phase locking experiment
- 1064-nm version

Frequency noise



Intensity noise



— Lowest free-running noise levels

- Stabilization by saturation signal of acetylene at 1542nm.
- Controllability

— High frequency noise @ high frequency

- Under investigations

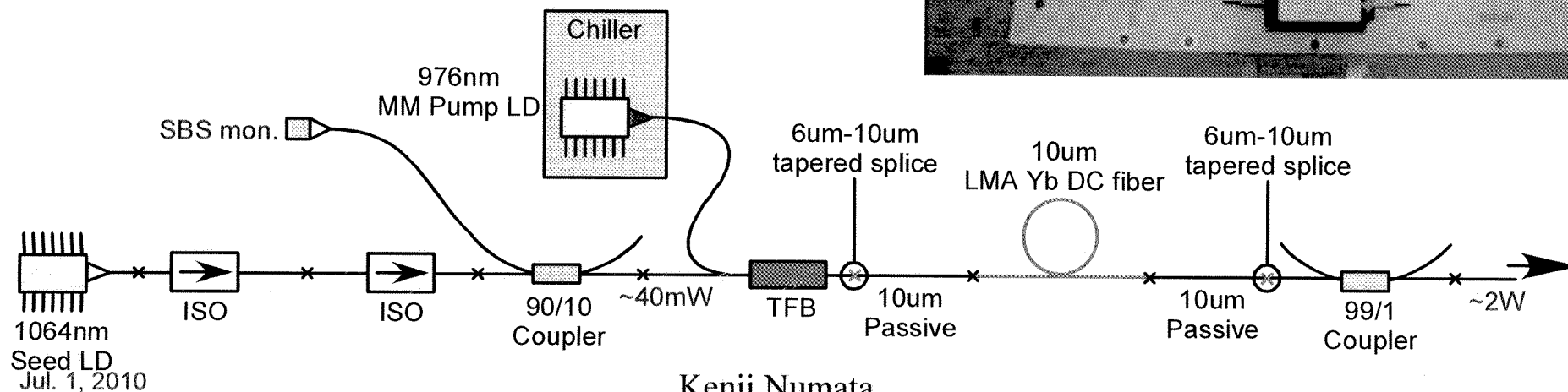
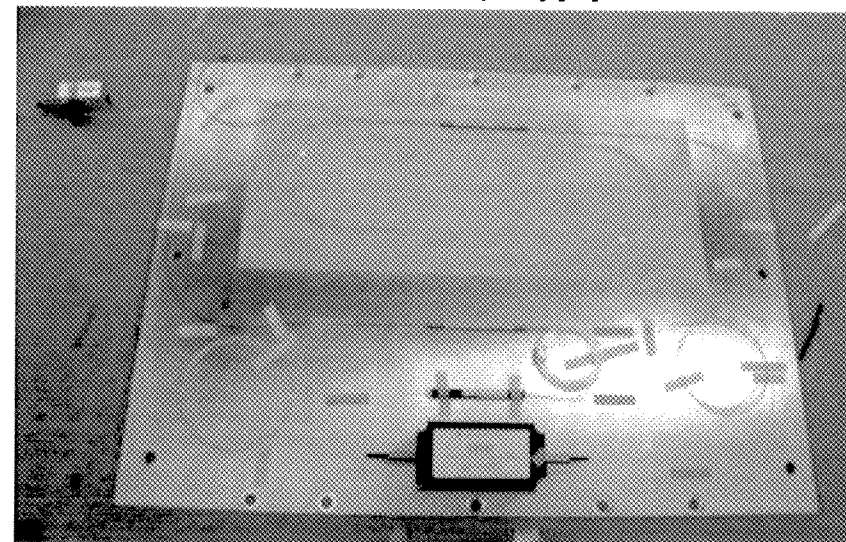
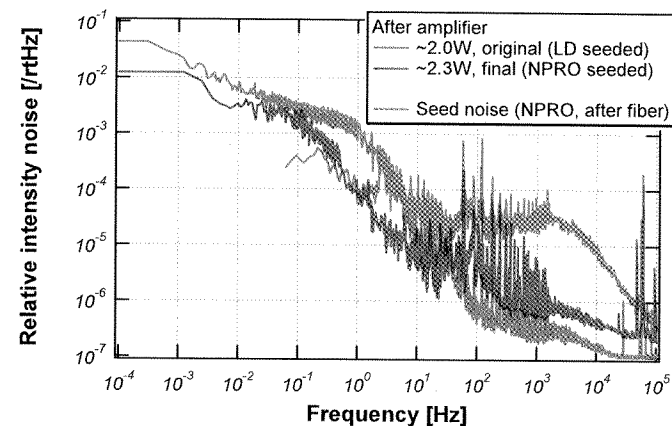
3. Fiber amplifier

Features

- Built by Lucent Government Solutions (LGS)
 - Clad pump, LMA fiber, ~4W maximum
- Focused on reliability
 - Detailed risk analysis
 - Passed thermal cycling tests

Status

- Noise measurements at GSFC
- Stabilization experiments



Kenji Numata

Space qualification tests

– To be done in collaboration with LGS

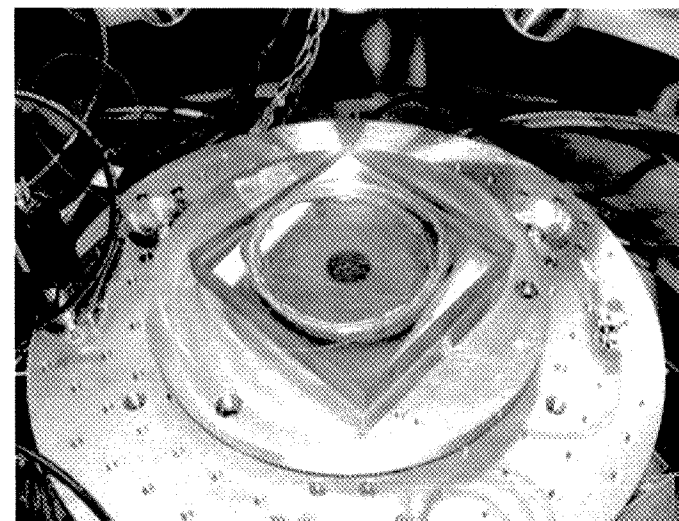
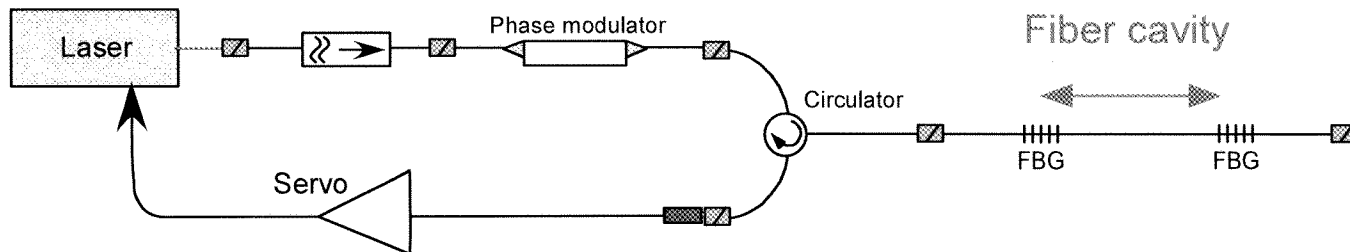
- Proton test (@ UC Davis)
- Gamma test (@ GSFC, 7/19~)
- Fiber components to be radiated
 - Fiber bragg grating (FBG), circulator (isolator), Band-pass filters, gain fibers, etc.
- Outgass, pyroshock



All-fiber frequency stabilization

– FBG Fabry-Perot cavity

- Finesse ~300



- ❏ Fiber approach very promising for space applications
 - Higher robustness, cleaner output, no strong magnet, etc.
 - Redundancy can be easily added.
 - New technology introduced frequently
 - No choice for solid-state amp for LISA-type CW, low-power applications
- ❏ Fiber-based lasers
 - At low frequency, NPRO is *not* the best anymore.
 - Custom-made fiber laser possible.
 - Possible issue is high frequency noise at higher frequency
 - Can be suppressed by fast frequency actuators (e.g. waveguide EOM)
- ❏ Current & future activities
 - Radiation tests
 - Full stabilization & metrology experiments